# **Physics of Planetary Systems — Exercises**

Astrophysikalisches Institut und Universitätssternwarte Jena Thüringer Landessternwarte Tautenburg

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## Set 2

### Problem 2.1

Figure 1 shows the correlation between period and semi-major axis of known planets as given in the Extrasolar Planet Encyclopaedia (exoplanet.eu). Why is there such a tight correlation? What about the outliers (some are circled)? Do these violate any physical law? What does this figure tell you? (2 points)



**Figure 1:** Correlation between period and semi-major axis. The left panel is in linear scale, the right one in log-log scale. (Source: exoplanet.eu)

#### Problem 2.2

Go to the Exoplanet Encyclopaedia (exoplanet.eu) and find all radial velocity detected extrasolar planets with masses  $1 \mathcal{M}_{Jup} < \mathcal{M} < 2 \mathcal{M}_{Jup}$ . (Click on "candidates detected by radial velocity". Under "planet table" you can sort by mass.) *Estimate* the probability that *all* of these planets are in reality companions with true mass greater or equal to that of a brown dwarf (i.e.  $\geq 20 \mathcal{M}_{Jup}$ ). (3 points)

#### Problem 2.3

Consider the following naïve model of the Sun's formation. An "infinitely" large and cold, nonrotating spherical cloud of molecular hydrogen with  $\mathcal{M} = \mathcal{M}_{\odot}$  collapses to the radius  $R = R_{\odot}$ , yielding the Sun. Which mechanism has stopped the collapse? Is the total energy conserved during the collapse? Estimate the temperature *T* of the cloud in its final state, i.e. the temperature of the Sun. Hint: use the virial theorem. (*3 points*)

#### Problem 2.4

Do you think planetary systems have better chances to develop around stars of lower or higher mass (or luminosity)? Why? (2 points)