

General Theory of Relativity

FSU Jena - WS 2009/2010

Problem set 07 - Solutions

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Problem 01 (Carroll, Problem 3.5)

Preconsiderations

Let M be the considered manifold with coordinate vector-basis $\{\partial_\mu\}$. A curve $x(t) : I \subseteq \mathbb{R} \rightarrow M$ is geodesic if

$$\nabla_{\dot{x}} \dot{x} = 0$$

where

$$\dot{x} := \frac{dx^\mu}{dt} \partial_\mu = \dot{x}^\mu \partial_\mu$$

is the tangent vector generated by the curve. Because of

$$\nabla_{\dot{x}} \dot{x} = \dot{x}^\mu \nabla_{\partial_\mu} (\dot{x}^\nu \partial_\nu) = \underbrace{\dot{x}^\mu (\partial_\mu \dot{x}^\nu)}_{\ddot{x}^\nu} \partial_\nu + \dot{x}^\mu \dot{x}^\lambda \nabla_{\partial_\mu} \partial_\lambda = \ddot{x}^\nu \partial_\nu + \dot{x}^\mu \dot{x}^\lambda \Gamma_{\mu\lambda}^\nu \partial_\nu$$

this is equivalent to

$$\ddot{x}^\nu + \Gamma_{\mu\lambda}^\nu \dot{x}^\mu \dot{x}^\lambda = 0 \tag{0.1}$$

A point-set $P \subseteq M$ is said to be geodesic, if it allows a smooth, geodesic parametrization $x(t)$. It turns out, that if P is indeed geodesic, any parametrization $x(t)$ with $g(\dot{x}, \dot{x}) = \text{const} \neq 0$ is geodesic.

The parallel transport of a vector $V \in T_p M$ along a curve $x : I \subseteq \mathbb{R} \rightarrow M$ is defined as the family of tangential vectors $(V(t))_{t \in I}$, $V(t) \in T_{x(t)} M$ for which

$$\nabla_{\dot{x}(t)} V(t) = 0$$

that is

$$0 = \dot{x}^\mu \nabla_{\partial_\mu} V^\nu \partial_\nu = \underbrace{\dot{x}^\mu (\partial_\mu V^\nu)}_{\dot{V}^\nu} \partial_\nu + \dot{x}^\mu V^\lambda \nabla_{\partial_\mu} \partial_\lambda = \dot{V}^\nu \partial_\nu + \Gamma_{\mu\lambda}^\nu \dot{x}^\mu V^\lambda \partial_\nu$$

or

$$\dot{V}^\nu + \Gamma_{\mu\lambda}^\nu \dot{x}^\mu V^\lambda = 0 \tag{0.2}$$

The Christoffel-Symbols are for this special metric given by

$$\begin{aligned} \Gamma_{\vartheta\vartheta}^\varphi &= 0 & \Gamma_{\varphi\varphi}^\vartheta &= -\sin \vartheta \cos \vartheta \\ \Gamma_{\vartheta\varphi}^\vartheta &= \Gamma_{\varphi\vartheta}^\vartheta = 0 & \Gamma_{\varphi\vartheta}^\varphi &= \Gamma_{\vartheta\varphi}^\varphi = \cot \vartheta \\ \Gamma_{\vartheta\vartheta}^\vartheta &= 0 & \Gamma_{\varphi\varphi}^\vartheta &= 0 \end{aligned}$$

(compare to problem set 06, Problem 02).

(a) Consider the curve $x(t) : I \subseteq \mathbb{R} \rightarrow S^2$ of constant φ , that is

$$x(t) := (\vartheta(t), \varphi(t)) := (\vartheta_0 + v_\vartheta \cdot t, \varphi_0) \quad , \quad \vartheta_0, v_\vartheta, \varphi_0 : \text{const}$$

Then through

$$\ddot{\vartheta} + \Gamma_{\mu\nu}^\vartheta \dot{x}^\mu \dot{x}^\nu = \underbrace{\ddot{\vartheta}}_0 + \Gamma_{\varphi\varphi}^\vartheta \dot{\varphi} \dot{\varphi} = 0$$

$$\ddot{\varphi} + \Gamma_{\mu\nu}^\varphi \dot{x}^\mu \dot{x}^\nu = \underbrace{\ddot{\varphi}}_0 + 2\Gamma_{\vartheta\varphi}^\varphi \dot{\vartheta} \dot{\varphi} = 0$$

one sees that $x(t)$ is indeed geodesic.

Let w.l.o.g. $\vartheta_0 \notin \{0, \pi\}$. For the curve

$$x(t) := (\vartheta(t), \varphi(t)) := (\vartheta_0, \varphi_0 + v_\varphi \cdot t) \quad , \quad \vartheta_0, \varphi_0, v_\varphi : \text{const}, \quad v_\varphi \neq 0$$

$$g(\dot{x}, \dot{x}) = g_{\varphi\varphi} v_\varphi^2 = \sin^2 \vartheta_0 \cdot v_\varphi^2 : \text{const} \neq 0$$

it holds

$$\ddot{\vartheta} + \Gamma_{\mu\nu}^\vartheta \dot{x}^\mu \dot{x}^\nu = \underbrace{\ddot{\vartheta}}_0 + \Gamma_{\varphi\varphi}^\vartheta \dot{\varphi} \dot{\varphi} = -\sin \vartheta \cos \vartheta \cdot v_\varphi^2$$

$$\ddot{\varphi} + \Gamma_{\mu\nu}^\varphi \dot{x}^\mu \dot{x}^\nu = \underbrace{\ddot{\varphi}}_0 + 2\Gamma_{\vartheta\varphi}^\varphi \dot{\vartheta} \dot{\varphi} = 0$$

from which follows that, $x(t)$ is geodesic if and only if $\sin \vartheta \cos \vartheta = 0$, that is $\cos \vartheta = 0$ and thus $\vartheta = \pi/2$.

(b) Consider the curve (circle) of constant latitude

$$x(t) := (\vartheta(t), \varphi(t)) := (\vartheta_0, t) \quad , \quad t \in [0, 2\pi]$$

and the vector $V = (1, 0)$ at $(\vartheta_0, \varphi = 0)$. Then, differential equation (0.2) takes the form

$$0 = \dot{V}^\vartheta + \Gamma_{\mu\nu}^\vartheta \dot{x}^\mu V^\nu = \dot{V}^\vartheta + \Gamma_{\varphi\varphi}^\vartheta \underbrace{\dot{\varphi}}_1 V^\varphi = \dot{V}^\vartheta - V^\varphi \sin \vartheta_0 \cos \vartheta_0$$

$$0 = \dot{V}^\varphi + \Gamma_{\mu\nu}^\varphi \dot{x}^\mu V^\nu = \dot{V}^\varphi + \Gamma_{\varphi\vartheta}^\varphi \underbrace{\dot{\varphi}}_1 V^\vartheta + \Gamma_{\vartheta\varphi}^\varphi \underbrace{\dot{\vartheta}}_0 V^\varphi = \dot{V}^\varphi + V^\vartheta \cot \vartheta_0$$

Differentiating the first equation we obtain the initial-value-problem (IVP)

$$\ddot{V}^\vartheta = -V^\vartheta \cos^2 \vartheta_0$$

whereas

$$V^\vartheta(0) = 1 \quad , \quad \dot{V}^\vartheta(0) = V^\varphi(0) \sin \vartheta_0 \cos \vartheta_0 = 0$$

and thus

$$V^\vartheta(t) = \cos [t \cdot \cos \vartheta_0] \tag{0.3}$$

Using (0.3) we obtain the IVP

$$\dot{V}^\varphi = -V^\vartheta \cot \vartheta_0 = -\cos [t \cdot \cos \vartheta_0] \cdot \cot \vartheta_0 \quad , \quad V^\varphi(0) = 0$$

with solution

$$V^\varphi(t) = -\frac{\sin [t \cdot \cos \vartheta_0]}{\sin \vartheta_0} \tag{0.4}$$

Thus the transport of $V = (1, 0)$ along $x(t)$ is given by

$$V(t) = (V^\vartheta(t), V^\varphi(t)) = \left[\cos [t \cdot \cos \vartheta_0], -\frac{\sin [t \cdot \cos \vartheta_0]}{\sin \vartheta_0} \right] \quad (0.5)$$

The transport along one circle ($t = 2\pi$) results in the vector

$$V_{\text{circle}} = (V_{\text{circle}}^\vartheta, V_{\text{circle}}^\varphi) = \left[\cos [2\pi \cos \vartheta_0], -\frac{\sin [2\pi \cos \vartheta_0]}{\sin \vartheta_0} \right] \quad (0.6)$$

Problem 02 (Carroll, Problem 3.6)

(a) We shall parametrize the space-time-paths of the two clocks by their respective proper-times τ_1, τ_2

$$x_i(t) := (t_i, r_i, \vartheta_i, \varphi_i)(\tau_i) := (t_i(\tau_i), R_i, \vartheta_i, \varphi_i) \quad , \quad \vartheta_i, \varphi_i, R_i : \text{const}, \quad i = 1, 2$$

(w.l.o.g. $\dot{t}_i \geq 0$) so that

$$1 = |g(\dot{x}_i, \dot{x}_i)| = |g_{tt}\dot{t}_i| = |(1 + 2\Phi(R_i))| \cdot \dot{t}_i$$

and thus

$$\frac{d\tau_i}{dt} = \frac{d\tau_i}{dt_i} = (\dot{t}_i)^{-1} = |1 + 2\Phi(R_i)|^{|\Phi| \leq 1/2} 1 - \frac{2GM}{R_i}$$

Hence, the time elapsed τ on each clock as a function of the coordinate time t is given by

$$\tau_i(t) = \left(1 - \frac{2GM}{R_i} \right) \cdot t \quad (0.7)$$

In particular, clock 2 (further away from earth's center) runs faster.

(b) Let

$$x(\tau) := (t(\tau), r(\tau), \vartheta(\tau), \varphi(\tau)) := (t(\tau), r_0, \vartheta_0, \varphi(\tau)) \quad , \quad r_0, \vartheta_0 := \frac{\pi}{2} : \text{const} \quad (0.8)$$

be the sought geodesic, whereas w.l.o.g. $g(\dot{x}, \dot{x}) = -1$ (parametrized by proper-time). Then $x(\tau)$ solves the Euler-Lagrange equations

$$\frac{d}{d\tau} \frac{\partial L}{\partial \dot{x}^\mu} - \frac{\partial L}{\partial x^\mu} = 0$$

for the Lagrangian

$$L(x, \dot{x}) := g(\dot{x}, \dot{x}) = g_{tt}\dot{t}^2 + g_{rr}\dot{r}^2 + g_{\vartheta\vartheta}\dot{\vartheta}^2 + g_{\varphi\varphi}\dot{\varphi}^2$$

Specifically,

$$\begin{aligned} 0 &= \frac{d}{d\tau} \frac{\partial L}{\partial \dot{r}} - \frac{\partial L}{\partial r} = \frac{d}{d\tau} \underbrace{(2\dot{r}g_{rr})}_{2\ddot{r}g_{rr} + 2\dot{r}\partial_\tau g_{rr}=0} - \dot{t}^2 \partial_r g_{tt} - \underbrace{\dot{r}^2}_{0} \partial_r g_{rr} - \underbrace{\dot{\vartheta}^2}_{0} \partial_r g_{\vartheta\vartheta} - \dot{\varphi}^2 \partial_r g_{\varphi\varphi} \\ &= \dot{t}^2 \cdot \frac{2GM}{r_0^2} - \dot{\varphi}^2 \cdot 2r_0 \underbrace{\sin^2 \vartheta_0}_1 \end{aligned}$$

is equivalent to

$$\left(\frac{d\varphi}{dt} \right)^2 = \frac{\dot{\varphi}^2}{\dot{t}^2} = \frac{GM}{r_0^3}$$

(w.l.o.g. $\frac{d\varphi}{dt} \geq 0$) or equivalently

$$\frac{d\varphi}{dt} = \sqrt{\frac{GM}{r_0^3}} \quad (0.9)$$

Moreover,

$$0 = \frac{d}{d\tau} \frac{\partial L}{\partial \dot{t}} - \frac{\partial L}{\partial t} = \frac{d}{d\tau} (2\dot{t}g_{tt}) = 2\ddot{t} \underbrace{g_{tt}}_{\text{const}}$$

is equivalent to $\ddot{t} = 0$, that is,

$$\boxed{t(\tau) = t_0 + v_t \cdot \tau, \quad t_0, v_t : \text{const}} \quad (0.10)$$

Similarly,

$$0 = \frac{d}{d\tau} \frac{\partial L}{\partial \dot{\varphi}} - \frac{\partial L}{\partial \varphi} = \frac{d}{d\tau} (2\dot{\varphi}g_{\varphi\varphi}) = 2\ddot{\varphi}g_{\varphi\varphi}$$

is equivalent to $\ddot{\varphi} = 0$, that is,

$$\boxed{\varphi(\tau) = \varphi_0 + v_\varphi \cdot \tau, \quad t_0 : \text{const}, \quad v_\varphi \stackrel{(0.9)}{=} v_t \cdot \sqrt{\frac{GM}{r_0^3}}} \quad (0.11)$$

Finally, assumption $g(\dot{x}, \dot{x}) = -1$, that is

$$-1 = g_{tt}\dot{t}^2 + g_{\varphi\varphi}\dot{\varphi}^2 = \left(\frac{2GM}{r_0} - 1\right) \cdot v_t^2 + r_0^2 \underbrace{\sin^2 \vartheta_0}_1 \cdot v_\varphi^2$$

leads to

$$\boxed{v_t = \sqrt{\frac{1}{1 - \frac{3GM}{r_0}}}} \quad (0.12)$$

Concluding, the geodesic at radius r_0 is given by

$$\boxed{(t(\tau), r(\tau), \vartheta(\tau), \varphi(\tau)) = \left[t_0 + \sqrt{\frac{1}{1 - \frac{3GM}{r_0}}} \cdot \tau, \quad r_0, \quad \frac{\pi}{2}, \quad \varphi_0 + \frac{1}{r_0} \sqrt{\frac{GM}{r_0 - 3GM}} \cdot \tau \right]} \quad (0.13)$$

Note: Because of

$$0 = \frac{d}{d\tau} \frac{\partial L}{\partial \dot{\vartheta}} - \frac{\partial L}{\partial \vartheta} = \frac{d}{d\tau} (2\dot{\vartheta}g_{\vartheta\vartheta}) - \dot{\varphi}^2 \partial_\vartheta g_{\varphi\varphi} = 2 \underbrace{\ddot{\vartheta}}_0 g_{\vartheta\vartheta} + 2 \underbrace{\dot{\vartheta}}_0 \frac{dg_{\vartheta\vartheta}}{d\tau} - \dot{\varphi}^2 \cdot 2r_0^2 \sin \vartheta_0 \underbrace{\cos \vartheta_0}_0 = 0$$

curve (0.13) is indeed geodesic.

Problem 03

Help-statement

A vector-field \mathbf{K} is Killing if and only if it satisfies

$$g(\nabla_{\mathbf{X}}\mathbf{K}, \mathbf{Y}) + g(\nabla_{\mathbf{Y}}\mathbf{K}, \mathbf{X}) = 0 \quad (0.14)$$

for any vector fields $\mathbf{X} = X^\mu \partial_\mu$, $\mathbf{Y} = Y^\mu \partial_\mu$.

Proof: Condition $\nabla_{(\nu} K_{\mu)} = 0$, that is,

$$\begin{aligned} 0 &= \underbrace{(\nabla_{\partial_\nu} g(\mathbf{K}, \cdot))_\mu}_{\substack{\nabla_{\partial_\nu}(g(\mathbf{K}, \partial_\mu)) \\ -g(\mathbf{K}, \nabla_{\partial_\nu} \partial_\mu)}} + \underbrace{(\nabla_{\partial_\mu} g(\mathbf{K}, \cdot))_\nu}_{\substack{\nabla_{\partial_\mu}(g(\mathbf{K}, \partial_\nu)) \\ -g(\mathbf{K}, \nabla_{\partial_\mu} \partial_\nu)}} \\ &= \left[g(\mathbf{K}, \nabla_{\partial_\nu} \partial_\mu) + g(\nabla_{\partial_\nu} \mathbf{K}, \partial_\mu) \right] - g(\mathbf{K}, \nabla_{\partial_\nu} \partial_\mu) + \left[g(\nabla_{\partial_\mu} \mathbf{K}, \partial_\nu) + g(\mathbf{K}, \nabla_{\partial_\mu} \partial_\nu) \right] - g(\mathbf{K}, \nabla_{\partial_\mu} \partial_\nu) \\ &= g(\nabla_{\partial_\nu} \mathbf{K}, \partial_\mu) + g(\nabla_{\partial_\mu} \mathbf{K}, \partial_\nu) \end{aligned}$$

is equivalent to

$$\begin{aligned} 0 &= g(\nabla_{\partial_\nu} \mathbf{K}, \partial_\mu) X^\nu Y^\mu + g(\nabla_{\partial_\mu} \mathbf{K}, \partial_\nu) X^\nu Y^\mu \\ &= g(\nabla_{\mathbf{X}} \mathbf{K}, \mathbf{Y}) + g(\nabla_{\mathbf{Y}} \mathbf{K}, \mathbf{X}) \quad \forall \text{ vectorfields } \mathbf{X}, \mathbf{Y} \end{aligned}$$

Note that use has been made of the metric-compatibility of the Levi-Civita-connection

$$\nabla_{\mathbf{X}}(g(\mathbf{Y}, \mathbf{Z})) = g(\nabla_{\mathbf{X}} \mathbf{Y}, \mathbf{Z}) + g(\mathbf{Y}, \nabla_{\mathbf{X}} \mathbf{Z}) \quad \forall \mathbf{X}, \mathbf{Y}, \mathbf{Z}$$

□

Proof of main statement

To be shown is $\mathbf{p}g(\mathbf{K}, \mathbf{p}) = 0$. But indeed, it follows from metric-compatibility of the Levi-Civita-connection

$$\nabla_{\mathbf{X}}(g(\mathbf{Y}, \mathbf{Z})) = g(\nabla_{\mathbf{X}} \mathbf{Y}, \mathbf{Z}) + g(\mathbf{Y}, \nabla_{\mathbf{X}} \mathbf{Z}) \quad \forall \mathbf{X}, \mathbf{Y}, \mathbf{Z}$$

that

$$\underbrace{\mathbf{p}g(\mathbf{K}, \mathbf{p})}_{\nabla_{\mathbf{p}}g(\mathbf{K}, \mathbf{p})} = \underbrace{g(\nabla_{\mathbf{p}} \mathbf{K}, \mathbf{p})}_0 + g(\mathbf{K}, \underbrace{\nabla_{\mathbf{p}} \mathbf{p}}_0) = 0$$

Note that use has been made of the help-statement to show

$$0 \stackrel{(0.14)}{=} g(\nabla_{\mathbf{p}} \mathbf{K}, \mathbf{p}) + g(\nabla_{\mathbf{p}} \mathbf{K}, \mathbf{p}) = 2g(\nabla_{\mathbf{p}} \mathbf{K}, \mathbf{p})$$

□

Problem 04

Preconsiderations

As the equation to be shown

$$\nabla_\alpha \nabla_\beta K^\mu = R_{\beta\alpha\nu}^\mu K^\nu$$

is a tensorial one, it suffices to show it's validity in any arbitrary coordinates. As is known, on any given point $p \in M$, there exists a local chart such that

$$(g_{\mu\nu})|_p = \text{diag}(\epsilon_1, \dots, \epsilon_n), \quad \epsilon_i \in \{\pm 1\}$$

and $\partial_\alpha g_{\mu\nu}|_p = 0$ (Riemann-normal coordinates).

As \mathbf{K} is Killing, that is

$$\begin{aligned} 0 &= \partial_\mu K_\nu + \partial_\nu K_\mu - 2\Gamma_{\nu\mu}^\lambda K_\lambda = \partial_\mu(K^\varkappa g_{\varkappa\nu}) + \partial_\nu(K^\varkappa g_{\varkappa\mu}) - 2\Gamma_{\nu\mu}^\lambda g_{\lambda\varkappa} K^\varkappa \\ &= K^\varkappa \underbrace{[\partial_\mu g_{\varkappa\nu} + \partial_\nu g_{\varkappa\mu}]}_{2\Gamma_{\mu\nu}^\lambda g_{\lambda\varkappa} + \partial_\varkappa g_{\mu\nu}} + g_{\varkappa\nu} \partial_\mu K^\varkappa + g_{\varkappa\mu} \partial_\nu K^\varkappa - 2\Gamma_{\nu\mu}^\lambda g_{\lambda\varkappa} K^\varkappa \\ &= K^\varkappa \partial_\varkappa g_{\mu\nu} + g_{\varkappa\nu} \partial_\mu K^\varkappa + g_{\varkappa\mu} \partial_\nu K^\varkappa \end{aligned}$$

it also fulfills

$$0 = \partial_\sigma (K^\varkappa \partial_\varkappa g_{\mu\nu} + g_{\varkappa\nu} \partial_\mu K^\varkappa + g_{\varkappa\mu} \partial_\nu K^\varkappa) \stackrel{\text{normal coordinates}}{=} K^\varkappa \partial_{\sigma\varkappa} g_{\mu\nu} + g_{\varkappa\nu} \partial_{\sigma\mu} K^\varkappa + g_{\varkappa\mu} \partial_{\sigma\nu} K^\varkappa \quad (0.15)$$

With

$$\begin{aligned}
\nabla_\alpha \nabla_\beta K^\mu &= \partial_\alpha (\nabla_\beta K^\mu) + \Gamma_{\alpha\lambda}^\mu \nabla_\beta K^\lambda - \Gamma_{\alpha\beta}^\lambda \nabla_\lambda K^\mu \\
&= \partial_\alpha (\partial_\beta K^\mu + \Gamma_{\beta\kappa}^\mu K^\kappa) + \Gamma_{\alpha\lambda}^\mu (\partial_\beta K^\lambda + \Gamma_{\beta\kappa}^\lambda K^\kappa) - \Gamma_{\alpha\beta}^\lambda (\partial_\lambda K^\mu + \Gamma_{\lambda\kappa}^\mu K^\kappa) \\
&= \partial_{\alpha\beta} K^\mu + (\partial_\alpha \Gamma_{\beta\kappa}^\mu) K^\kappa + \Gamma_{\beta\kappa}^\mu \partial_\alpha K^\kappa + \Gamma_{\alpha\kappa}^\mu \partial_\beta K^\kappa + \Gamma_{\alpha\lambda}^\mu \Gamma_{\beta\kappa}^\lambda K^\kappa - \Gamma_{\alpha\beta}^\lambda \partial_\lambda K^\mu - \Gamma_{\alpha\beta}^\lambda \Gamma_{\lambda\kappa}^\mu K^\kappa \\
&\stackrel{\text{normal coordinates}}{=} \partial_{\alpha\beta} K^\mu + (\partial_\alpha \Gamma_{\beta\kappa}^\mu) K^\kappa
\end{aligned}$$

and

$$R^\mu{}_{\beta\alpha\kappa} K^\kappa \stackrel{\text{normal coordinates}}{=} (\partial_\alpha \Gamma_{\beta\kappa}^\mu) K^\kappa - (\partial_\kappa \Gamma_{\beta\alpha}^\mu) K^\kappa$$

it suffices to show

$$\partial_{\alpha\beta} K^\mu = -(\partial_\kappa \Gamma_{\beta\alpha}^\mu) K^\kappa \tag{0.16}$$

Final proof

Indeed,

$$\begin{aligned}
& -(\partial_\kappa \Gamma_{\beta\alpha}^\mu) K^\kappa \stackrel{\text{normal coordinates}}{=} -\frac{g^{\mu\lambda}}{2} (\partial_{\kappa\beta} g_{\alpha\lambda} + \partial_{\kappa\alpha} g_{\beta\lambda} - \partial_{\kappa\lambda} g_{\beta\alpha}) K^\kappa \\
& \stackrel{(0.15)}{=} -\frac{g^{\mu\lambda}}{2} (-g_{\kappa\lambda} \partial_{\beta\alpha} K^\kappa - g_{\kappa\alpha} \partial_{\beta\lambda} K^\kappa - g_{\kappa\lambda} \partial_{\alpha\beta} K^\kappa - g_{\kappa\beta} \partial_{\alpha\lambda} K^\kappa + g_{\kappa\alpha} \partial_{\lambda\beta} K^\kappa + g_{\kappa\beta} \partial_{\lambda\alpha} K^\kappa) \\
& = g^{\mu\lambda} g_{\kappa\lambda} \partial_{\alpha\beta} K^\kappa = \partial_{\alpha\beta} K^\mu
\end{aligned}$$

□