

# General Theory of Relativity

## FSU Jena - WS 2009/2010

### Problem set 06 - Solutions

Stilianos Louca

February 10, 2010

#### Problem 01

Defining *angles* between vectors  $X, Y$  as

$$X \angle Y := \frac{g(X, Y)}{\sqrt{g(X, X)} \cdot \sqrt{g(Y, Y)}}$$

(compare to definition in Euclidian & unitary spaces), leads to the preservation of angles by any conformal metric-transformation  $g_{\mu\nu}|_x = \Omega(x) \cdot \tilde{g}_{\mu\nu}|_x$ , as

$$\frac{g(X, Y)}{\sqrt{g(X, X)} \cdot \sqrt{g(Y, Y)}} = \frac{\Omega \cdot \tilde{g}(X, Y)}{\sqrt{\Omega \cdot \tilde{g}(X, X)} \cdot \sqrt{\Omega \cdot \tilde{g}(Y, Y)}}$$

Furthermore, for any vector  $X$ ,  $\tilde{g}(X, X) = 0$  implies  $\underbrace{g(X, X)}_{\Omega \cdot \tilde{g}(X, X)} = 0$ , that is, null-curves are still null-curves after conformal metric-transformations.

#### Problem 02

Starting with

$$\Gamma_{\mu\nu}^\lambda = \frac{1}{2} g^{\lambda\kappa} (\partial_\mu g_{\nu\kappa} + \partial_\nu g_{\mu\kappa} - \partial_\kappa g_{\mu\nu})$$

and a diagonal metric  $g_{\mu\nu}$ , we consider following cases:

1.  $\mu \neq \nu \neq \lambda \neq \mu$ : Then  $g_{\mu\nu} = 0$  and  $g^{\lambda\kappa} = 0$  for  $\lambda \neq \kappa$ , thus

$$\Gamma_{\mu\nu}^\lambda = \frac{1}{2} g^{\lambda\lambda} \left( \partial_\mu \underbrace{g_{\nu\lambda}}_0 + \partial_\nu \underbrace{g_{\mu\lambda}}_0 \right) = 0$$

2.  $\mu = \nu \neq \lambda$ , then

$$\Gamma_{\mu\mu}^\lambda = \frac{1}{2} \underbrace{g^{\lambda\kappa}}_{\substack{0 \\ \text{for} \\ \kappa \neq \lambda}} (2\partial_\mu g_{\mu\kappa} - \partial_\kappa g_{\mu\mu}) = \frac{1}{2} g^{\lambda\lambda} \left( 2\partial_\mu \underbrace{g_{\mu\lambda}}_0 - \partial_\lambda g_{\mu\mu} \right) = -\frac{1}{2} g^{\lambda\lambda} \partial_\lambda g_{\mu\mu} = -\frac{1}{2} \frac{1}{g_{\lambda\lambda}} \partial_\lambda g_{\mu\mu}$$

Note that the inverse  $g^{\mu\nu}$  of the diagonal  $g_{\mu\nu}$  is the diagonal matrix with entries  $g^{\mu\mu} = 1/g_{\mu\mu}$ .

3.  $\nu = \lambda \neq \mu$ , then

$$\Gamma_{\mu\lambda}^\lambda = \frac{1}{2} \underbrace{g^{\lambda\kappa}}_{\substack{0 \\ \text{for} \\ \kappa \neq \lambda}} (\partial_\mu g_{\lambda\kappa} + \partial_\lambda g_{\mu\kappa} - \partial_\kappa g_{\mu\lambda}) = \frac{1}{2} g^{\lambda\lambda} (\partial_\mu g_{\lambda\lambda} + \cancel{\partial_\lambda g_{\mu\lambda}} - \cancel{\partial_\lambda g_{\mu\lambda}}) = \frac{1}{2} \underbrace{g^{\lambda\lambda}}_{1/g_{\lambda\lambda}} \partial_\mu g_{\lambda\lambda}$$

$$\stackrel{(\clubsuit)}{=} \frac{1}{2} \frac{1}{|g_{\lambda\lambda}|} \partial_\mu |g_{\lambda\lambda}| = \partial_\mu \left( \ln \sqrt{|g_{\lambda\lambda}|} \right)$$

(♣): Note that, as  $g$  is invertible and diagonal, all entries  $g_{\lambda\lambda}$  are non-zero. As  $g$  is further smooth on  $M$ , it's entries are also non-zero with constant signum in some open neighborhood of the considered point, thus

$$\begin{aligned}\frac{\partial_\mu |g_{\lambda\lambda}|(\mathbf{x})}{|g_{\lambda\lambda}|(\mathbf{x})} &= \frac{1}{|g_{\lambda\lambda}(\mathbf{x})|} \cdot \lim_{h \rightarrow 0} \frac{|g_{\lambda\lambda}(\mathbf{x} + h\mathbf{e}_\mu)| - |g_{\lambda\lambda}(\mathbf{x})|}{h} \\ &= \frac{\text{sgn}(g_{\lambda\lambda}(\mathbf{x}))}{\text{sgn}(g_{\lambda\lambda}(\mathbf{x}))} \cdot \frac{1}{g_{\lambda\lambda}(\mathbf{x})} \cdot \lim_{h \rightarrow 0} \frac{g_{\lambda\lambda}(\mathbf{x} + h\mathbf{e}_\mu) - g_{\lambda\lambda}(\mathbf{x})}{h} \\ &= \frac{\partial_\mu g_{\lambda\lambda}}{g_{\lambda\lambda}}\end{aligned}$$

4.  $\mu = \nu = \lambda$ , then

$$\Gamma_{\lambda\lambda}^\lambda = \frac{1}{2} \underbrace{g^{\lambda\lambda}}_{\substack{0 \\ \text{for} \\ \lambda \neq \lambda}} (\partial_\lambda g_{\lambda\lambda} + \partial_\lambda g_{\lambda\lambda} - \partial_\lambda g_{\lambda\lambda}) = \frac{1}{2} g^{\lambda\lambda} \partial_\lambda g_{\lambda\lambda} = \frac{1}{2} \frac{\partial_\lambda g_{\lambda\lambda}}{g_{\lambda\lambda}} \stackrel{\text{similarly to (3)}}{=} \partial_\lambda \ln \sqrt{|g_{\lambda\lambda}|}$$

### Problem 03

Beginning with the Euklidian metric in spherical coordinates

$$g_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & r^2 & 0 \\ 0 & 0 & r^2 \sin^2 \vartheta \end{pmatrix}$$

and using the results from (02) we write

$$\Gamma_{r\vartheta}^\varphi = \Gamma_{\vartheta r}^\varphi = \Gamma_{r\varphi}^\vartheta = \Gamma_{\varphi r}^\vartheta = \Gamma_{\vartheta\varphi}^r = \Gamma_{\varphi\vartheta}^r = 0$$

$$\Gamma_{rr}^\vartheta = -\frac{1}{2} \frac{1}{g_{\vartheta\vartheta}} \underbrace{\partial_\vartheta g_{rr}}_0 = 0$$

$$\Gamma_{rr}^\varphi = -\frac{1}{2} \frac{1}{g_{\varphi\varphi}} \underbrace{\partial_\varphi g_{rr}}_0 = 0$$

$$\Gamma_{\vartheta\vartheta}^r = -\frac{1}{2} \frac{1}{g_{rr}} \underbrace{\partial_r g_{\vartheta\vartheta}}_{2r} = -r$$

$$\Gamma_{\vartheta\vartheta}^\varphi = -\frac{1}{2} \frac{1}{g_{\varphi\varphi}} \underbrace{\partial_\varphi g_{\vartheta\vartheta}}_0 = 0$$

$$\Gamma_{\varphi\varphi}^r = -\frac{1}{2} \frac{1}{g_{rr}} \underbrace{\partial_r g_{\varphi\varphi}}_{2r \sin^2 \vartheta} = -r \sin^2 \vartheta$$

$$\Gamma_{\varphi\varphi}^\vartheta = -\frac{1}{2} \frac{1}{g_{\vartheta\vartheta}} \underbrace{\partial_\vartheta g_{\varphi\varphi}}_{2r^2 \sin \vartheta \cos \vartheta} = -\sin \vartheta \cos \vartheta$$

$$\Gamma_{r\vartheta}^\vartheta = \Gamma_{\vartheta r}^\vartheta = \frac{1}{2} \frac{1}{g_{\vartheta\vartheta}} \underbrace{\partial_r g_{\vartheta\vartheta}}_{2r} = \frac{1}{r}$$

$$\Gamma_{r\varphi}^\varphi = \Gamma_{\varphi r}^\varphi = \frac{1}{2} \frac{1}{g_{\varphi\varphi}} \underbrace{\partial_r g_{\varphi\varphi}}_{2r \sin^2 \vartheta} = \frac{1}{r}$$

$$\Gamma_{\vartheta r}^r = \Gamma_{r\vartheta}^r = \frac{1}{2} \frac{1}{g_{rr}} \underbrace{\partial_\vartheta g_{rr}}_0 = 0$$

$$\Gamma_{\vartheta\varphi}^\varphi = \frac{1}{2} \frac{1}{g_{\varphi\varphi}} \underbrace{\partial_\vartheta g_{\varphi\varphi}}_{2r^2 \sin \vartheta \cos \vartheta} = \cot \vartheta$$

$$\Gamma_{\varphi r}^r = \Gamma_{r\varphi}^r = \frac{1}{2} \frac{1}{g_{rr}} \underbrace{\partial_\varphi g_{rr}}_0 = 0$$

$$\Gamma_{\varphi\vartheta}^\vartheta = \Gamma_{\vartheta\varphi}^\vartheta = \frac{1}{2} \frac{1}{g_{\vartheta\vartheta}} \underbrace{\partial_\varphi g_{\vartheta\vartheta}}_0 = 0$$

and finally

$$\Gamma_{rr}^r = \frac{1}{2} \frac{\partial r g_{rr}}{g_{rr}} = 0 \quad \Gamma_{\vartheta\vartheta}^\vartheta = \frac{\partial_\vartheta g_{\vartheta\vartheta}}{g_{\vartheta\vartheta}} = 0 \quad \Gamma_{\varphi\varphi}^\varphi = \frac{\partial_\varphi g_{\varphi\varphi}}{g_{\varphi\varphi}} = 0$$

### Problem 04 (Carroll, Problem 3.2)

As is known, in coordinates for which all first derivatives of the components  $g_{\mu\nu}$  vanish (Riemann-normal-coordinates), partial and covariant derivatives are identical, that is

$$\nabla_\mu V^\nu = \partial_\mu V^\nu = (\nabla_{\partial_\mu} V)^\nu$$

In any other coordinate system, the covariant derivative is given by

$$\nabla_\mu V^\nu = \partial_\mu V^\nu + \Gamma_{\mu\kappa}^\nu V^\kappa$$

As the tensor field defined by the covariant derivative does not depend on the coordinates chosen (in other words:  $\nabla$  creates tensors), it may be used to express fields, defined through partial derivatives in Riemann-normal-coordinates, in other coordinates as well. Thus for example in spherical coordinates:

$$\begin{aligned} \vec{\nabla} \Phi &\stackrel{\text{cartesian}}{=} (\partial^\mu \Phi) \partial_\mu = g^{\mu\nu} (\partial_\nu \Phi) \partial_\mu \stackrel{\text{general}}{=} g^{\mu\nu} (\nabla_\nu \Phi) \partial_\mu \stackrel{\text{spherical}}{\stackrel{(03)}{=}} (\partial_r \Phi) \partial_r + \frac{1}{r^2} (\partial_\vartheta \Phi) \partial_\vartheta + \frac{1}{r^2 \sin^2 \vartheta} (\partial_\varphi \Phi) \partial_\varphi \\ \text{div } \mathbf{V} &\stackrel{\text{cartesian}}{=} \partial_\mu V^\mu = \nabla_\mu V^\mu \stackrel{\text{general}}{=} \partial_\mu V^\mu + \Gamma_{\mu\kappa}^\mu V^\kappa \stackrel{\text{spherical}}{\stackrel{(03)}{=}} \partial_r V^r + \partial_\vartheta V^\vartheta + \partial_\varphi V^\varphi + \frac{2}{r} V^r + \cot \vartheta V^\vartheta \\ \text{rot } \mathbf{V} &\stackrel{\text{cartesian}}{=} \partial_\mu V_\nu \epsilon^{\mu\nu\kappa} \partial_\kappa = \nabla_\mu V^\lambda g_{\lambda\nu} \epsilon^{\mu\nu\kappa} \partial_\kappa \stackrel{\text{general}}{=} (\partial_\mu V^\lambda + \Gamma_{\mu\kappa}^\lambda V^\kappa) g_{\lambda\nu} \underbrace{\text{sgn}(\mu\nu\kappa)}_{\substack{1/r^2 \sin \vartheta \\ \text{for spherical}}} \frac{1}{\sqrt{|\det g|}} \cdot \partial_\kappa \\ &\stackrel{\text{spherical}}{\stackrel{(03)}{=}} \frac{1}{\sin \vartheta} \cdot [\sin^2 \vartheta \cdot \partial_\vartheta V^\varphi - \partial_\varphi V^\vartheta + 2 \sin \vartheta \cos \vartheta \cdot V^\varphi] \cdot \partial_r \\ &\quad + \frac{1}{r^2 \sin \vartheta} \cdot [\partial_\varphi V^r - r^2 \sin^2 \vartheta \cdot \partial_r V^\varphi - 2r \sin^2 \vartheta \cdot V^\varphi] \cdot \partial_\vartheta \\ &\quad + \frac{1}{r^2 \sin \vartheta} \cdot [r^2 \partial_r V^\vartheta - \partial_\vartheta V^r + 2r V^\vartheta] \cdot \partial_\varphi \end{aligned}$$

for any vector-field

$$\mathbf{V} = V^r \partial_r + V^\vartheta \partial_\vartheta + V^\varphi \partial_\varphi$$

### Notes:

- (i) Typically the gradient of  $\Phi \in \mathcal{F}$  is defined as the 1-form  $d\Phi$  for which  $d\Phi(X) = X\Phi$  for all vectors  $X$ :

$$d\Phi = \partial_\mu \Phi dx^\mu \stackrel{\text{spherical}}{=} \partial_r \Phi dr + \partial_\vartheta \Phi d\vartheta + \partial_\varphi \Phi d\varphi$$

As the vectorial gradient  $\vec{\nabla} \Phi$  is defined as the vector field for which

$$g(\vec{\nabla} \Phi, \cdot) = d\Phi$$

it is given by

$$\vec{\nabla} \Phi = g^{\mu\nu} \partial_\mu \Phi \partial_\nu$$

(indices pulled up).

- (ii) Typically, one finds in classical physics textbooks usage of the normalized coordinate-vectors  $\mathbf{e}_\mu$  instead of the coordinate-vectors  $\partial_\mu$ . Normalizing the basis  $\mathbf{e}_\mu := \partial_\mu / \sqrt{g(\partial_\mu, \partial_\mu)}$  (non-coordinate!), that is,

$$\mathbf{e}_r := \partial_r, \quad \mathbf{e}_\vartheta := \frac{1}{r} \partial_\vartheta, \quad \mathbf{e}_\varphi := \frac{1}{r \sin \vartheta} \partial_\varphi$$

and thus

$$\mathbf{V} = V^\mu \partial_\mu = \underbrace{V^\mu \sqrt{g(\partial_\mu, \partial_\mu)}}_{\tilde{V}^\mu} \cdot \mathbf{e}_\mu = \tilde{V}^\mu \mathbf{e}_\mu$$

that is,

$$\tilde{V}^r := V^r, \quad \tilde{V}^\vartheta := r V^\vartheta, \quad \tilde{V}^\varphi := r \sin \vartheta \cdot V^\varphi$$

yields

$$\vec{\nabla} \Phi = (\partial_r \Phi) \cdot \mathbf{e}_r + \frac{1}{r} (\partial_\vartheta \Phi) \cdot \mathbf{e}_\vartheta + \frac{1}{r \sin \vartheta} (\partial_\varphi \Phi) \cdot \mathbf{e}_\varphi$$

$$\operatorname{div} \mathbf{V} = \frac{1}{r^2} \partial_r (r^2 \tilde{V}^r) + \frac{1}{r \sin \vartheta} \left[ \partial_\vartheta (\sin \vartheta \cdot \tilde{V}^\vartheta) + \partial_\varphi \tilde{V}^\varphi \right]$$

$$\operatorname{rot} \mathbf{V} = \frac{1}{r \sin \vartheta} \left[ \partial_\vartheta (\sin \vartheta \cdot \tilde{V}^\varphi) - \partial_\varphi \tilde{V}^\vartheta \right] \cdot \mathbf{e}_r + \frac{1}{r} \left[ \frac{1}{\sin \vartheta} \partial_\varphi \tilde{V}^r - \partial_r (r \tilde{V}^\varphi) \right] \cdot \mathbf{e}_\vartheta + \frac{1}{r} \left[ \partial_r (r \tilde{V}^\vartheta) - \partial_\vartheta \tilde{V}^r \right] \cdot \mathbf{e}_\varphi$$

Comparing the latest form with [1] verifies the correctness of the results.

## References

- [1] *Taschenbuch Mathematischer Formeln*, H.J. Bartsch  
Fachbuchverlag Leipzig, 2004